Semest	S Stage 3A/3 er 2 Examin	B ation, 2011
Name:		
Teacher:_		
Time allowed	for this paper	
Reading time bef	ore commencing work:	Ten minutes
Working time for	paper:	Three hours
Material requ	ired/recommended f	or this paper
To be provided b	y the supervisor	
This Question/An	swer Booklet	
Physics: Formulae	e, Constants and Data Sheet	(inside front cover of this Question/Answer Booklet)
To be provided b	y the cand idate	
Standard Items:	Pens, pencil, eraser, corre	ction fluid, ruler, highlighter
Special Items:	non-programmable calcul Council for this course, dr.	ators satisfying the conditions set by the Curriculum awing templates, drawing compass and a protractor
Importantno	te to candidates	
No other items m have any unauth	ay be taken into the examin orised notes or other items	nation room. It is your responsibility to ensure that you of a non-personal nature in the examination room. If yo

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Structure of this paper

Section	Number of questions available	Number of questions to be answered	Suggested working time (minutes)	Marks availa ble	Percentage of exam
Section One:	13	13	50	54	30
Short response			20	2.	20
Section Two:	0	0	00	00	50
Problem-solving	ð	ŏ	90	90	50
Section Three:	_	_			
Comprehension	2	2	40	36	20
			180	180	100

Instructions to candidates

- 1. The rules for the conduct of Western Australian external examinations are detailed in the Year 12 Information Handbook 2010. Sitting this examination implies that you agree to abide by these rules.
- 2. Write answers in this Question/Answer Booklet.
- 3. You must be careful to confine your responses to the specific questions asked and to follow any instructions that are specific to a particular question.
- 4. Working or reasoning should be clearly shown when calculating or estimating answers.
- 5. Spare pages are included at the end of this booklet. They can be used for planning your responses and/or as additional space if required to continue an answer.
 - Planning: If you use the spare pages for planning, indicate this clearly at the top of the page.
 - Continuing an answer. If you need to use the space to continue an answer, indicate in the original answer space where the answer is continued, i.e. give the page number.
 Fill in the number of the question(s) that you are continuing to answer at the top of the page.

3 SECTION ONE: Short Response 54 marks (30%) This section has 13 questions. Answer ALL questions. Write your answers in the spaces provided. Suggested working time for this section is 54 minutes. Question 1 (2 marks) The diagram below shows a boat taking shelter from large waves behind a stone barrier. Explain why the boat is still at risk of damage from the waves. WAVES BARRIER

The wavelength of the waves is of similar size as the barrier, so significant diffraction around the barrier will occur and the boat will be buffeted by the diffracted waves.

Question 2

Hamish, who has a mass of 65.0 kg, is walking across a high wire suspended between two buildings. The wire makes an angle of 10° to the horizontal where it is attached at either end. Find the tension in the wire.

> 2Tsin10°= mg T= (65)9.8)/2sin10 = 1830 N





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Question 3

(4 marks)

The table below shows the 6 types of sub-atomic particles known as quarks and their properties:

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NAME	SYMBOL	Charge (Q)	Baryon Number (B)	Strangeness (S)	Charm (c)	Bottomness (b)	Topness (t)
Uр	u	$+\frac{2}{3}e$	$\frac{1}{3}$	0	0	0	0
Down	d	$-\frac{1}{3}e$	$\frac{1}{3}$	0	0	0	0
Strange	s	$-\frac{1}{3}e$	$\frac{1}{3}$	-1	0	0	0
Charmed	с	$+\frac{2}{3}e$	1 3	0	+1	0	0
Bottom	b	$-\frac{1}{3}e$	$\frac{1}{3}$	0	0	-1	0
Тор	t	$+\frac{2}{3}e$	1 3	0	0	0	+1

(a) Use the information in the table to complete the table below, to show the composition of the *Pion-minus* meson and the *Kaon-plus* meson.

NAME	SYMBOL	B	S	С	b	It	Quarks
Pion-plus	π^+	0	0	0	0	0	ud
Pion-minus	π ⁻	0	0	0	0	0	u d
Kaon-plus	K⁺	0	+1	0	0	0	น ร

(2 marks)

(b) The constituents of a proton can be represented by: uud and an antiproton by uud U se this notation to show the constituents of the following:

İ.	Neutron:	udd	

ii. Antineutron: udd

(2 marks)





Question 8

The mass of Pluto was not known until it was discovered that Pluto had a moon. Explain why this fact helped scientists determine Pluto's mass. Be sure to state what observations the scientists needed to make and include appropriate equations in your explanation.

The scientists needed to observe/measure the orbital radius R and orbital period T of Pluto's moon. With these values they could calculate the mass M of Pluto, as Kepler's third states:

$$M = \frac{4\pi^2 R^3}{GT^2}$$

Question 9

(6 marks)

(5 marks)

The diagrams below show different arrangements of current carrying conductors. Indicate clearly the shape and direction of the magnetic fields in each situation. In (b) be sure to indicate the shape of the field on the card.





b) How does Sophie explain that the two flashes of light reach Eloise simultaneously? (3 marks)

From Sophie's point of view, the front strike was first, then the back strike, which explains why Eloise (travelling to the left) sees them as simultaneous. i.e Sophie infers that Eloise was travelling to the left to meet the light from later strike, thus seeing it at the same time as the earlier strike.









(d) Scientists use the information in the two spectra as evidence that the Sun is spinning on its axis.

i. Which spectrum, **A** or **B** shows light that is emitted from the receding side of the Sun?

Spectrum B

(1 mark)

ii. Explain how the lines in the spectra show that one side of the Sun is spinning towards us (approaching), while the other side is spinning away from us (receding).

Spectrum A has its emission lines shifted toward the short wavelength end of the spectrum (blue shift). As the source of the light is approaching us, the waves are bunched up and therefore the wavelength shortened. The opposite is true for the light emiited from the receding edge of the Sun, so spectrum B is produced. (2 marks)

(e) Explain briefly how a line absorption spectrum is formed.

An absorption spectrum occurs when light passes through a cold, dilute gas and atoms in the gas absorb characteristic frequencies; since the re-emitted light is unlikely to be emitted in the same direction as the absorbed photon, this gives rise to dark lines (absence of light) in the spectrum.

(3 marks)







(c) At the instant shown in the diagram, the helicopter is 477 m behind the car. How long should the stuntman wait before he makes his jump? 6 marks

Needs to cover 191 m. At a relative speed of 40 m/s, this will take him:

t = s/v = 191/40 = 4.78 s

(d) What is the velocity of the stuntman, relative to the car when he lands in the car? (4 marks)



(e) Where is the helicopter when the stuntman lands in the car?

(2 marks)

The helicopter would be directly above the car as the stuntman and the helicopter have the same horizontal velocities.



19 (b) What power is lost in the 3.0 ohm transmission wires for each configuration? Configuration A: $P(lost) = I^2 R$ = 17.5²x 0.3 = 92 W Configuration B: $P(lost) = I^{2}R$ = 87.5²x 0.3 = 2300 W (4 marks) (c) By comparing the percentage efficiencies of the two configurations, make a recommendation to the farmer on which one to use. Configuration A has % loss of $(92/4200) \times 100 = 2.2\%$ Configuration B has a % loss of (2300/4200) x 100 = 54.7% Therefore configuration A is the best to use. (5 marks) SEE NEXT PAGE



Question 19

An engineering company is experimenting with a new type of transportation. It consists of a 365 kg aluminium glider that rests on two strong rails that provide an 825 Amp current that flows through the wings of the glider. A strong magnetic field in the region of the glider causes it to be propelled forward. The rails are 23.5 m apart.



The magnetic field is produced by powerful electromagnets.

- (a) Show the direction of this field on the diagram as dots (into the page) or crosses (out of he page)
- (b) Find the intensity of the magnetic field necessary to propel the glider forward with an acceleration of 2.6 ms⁻². You may ignore the effects of friction.

 $F = ma = 365 \times 2.6 = 949 N$ F = BIlso: B = F/Il $= 949/825 \times 23.5$ $= 4.89 \times 10 T$

(5 marks)

(1 mark)

(c) Gemma is monitoring the current flowing during the experiment and notices that it dropped as the glider gained speed. Explain her observation.

As the glider accelerates forwards, it cuts magnetic flux and generates an emf across its wings. The induced emf opposes the DC voltage supplied (back emf resulting from Lenz's Law). Hence the net voltage in the circuit decreases and so does the current flowing.

(3 marks)

(9 marks)



Question 21 (13 marks) Barnard's star is a red dwarf (ex-star) that has 17% of our Sun's mass and 15% of our Sun's diameter. Barnard's star still radiates heat (and light), but is much cooler than our own Sun. For Bemard's star to have a planet with liquid water at its surface, the planet would need to be quite close to the star's surface. According to astronomers, a suitable distance would be 0.06 AU. Such a planet is said to be in the "Habitable Zone". Note one AU (astronomical unit) and is the distance between the Earth and our Sun. (a) Determine the time, in weeks, that it would take a planet in the habitable zone to orbit Barnard's Star. M=(0.17) mass of Sun $T^2 = \frac{4\pi^2 R^3}{GM}$ R = 0.06 x Earth-Sun distance $T^{2} = \frac{4\pi^{2}(9 \times 10^{9})^{3}}{(6.67 \times 10^{-11})(3.4 \times 10^{29})}$ $T = 1.13 \times 10^6 s \sim 2$ weeks (5 marks) (b) Barnard's star is in the constellation Ophiuchus and has a parallax angle of 0.549 arc-second as measured from Earth. With the aid of a suitable diagram, explain what is meant by parallax angle and outline i how it is measured. Distant stars are used as a background to measure the angle subtended from a star to the line representing the diameter of the Earth's 7p orbit around the Sun (2 AU). This gives a simple Star in Question triangle where, tan(p) = 1 AU/d, where d is the distance to the star in question. If the angle is small tan(p) = p. So $d(in AU) = \frac{1AU/p}{(5 \text{ marks})}$ Background Stars



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(d)	Use your graph in (c) to determine (i) the total stopping distance <i>D</i> for a speed of 35 ms ⁻¹ .	
	At $v=35$, D/v is 3.35, so $D = 3.35 \times 35 = 117 \text{ m}$	[2 marks]
	(ii) the intercept on the D/v axis.	
	0.6 s	[1 mark]
	(iii) the gradient of the line of best fit.	
	$(3.35-1)/35-5 = 7.83 \times 10^{-2} s^2/m$	[2 marks]
(e)	Using your answers to (d)(ii) and (d)(iii), deduce the equation for D in the second se	terms of v.
	$D = 0.6v + 0.0783v^2$	
		[2 marks]
(f)	The uncertainty in the measurement of the distance D is \pm 0.3 m and t the measurement of the speed v is \pm 0.5 m s ⁻¹ .	he uncertainty in
	(i) For the data point corresponding to $v = 27.0 \text{ m s}^{-1}$, calculate the the value of $D_{V_{\rm L}}$	absolute uncertainty in
	2.74s +/- 0.062 s	
		[2 marks]
	(ii) Each of the data points in the table was obtained by taking the a values of <i>D</i> for each value of <i>v</i> . Suggest what effect, if any, the t have on the uncertainties in the data points.	verage of several aking of averages will
	Taking averages will reduce random errror in measureme	ent and hence reduce
	the uncertainties in data points.	
		[2 marks]

Question 23 WHAT IS THE UNIVERSE MADE OF?

(18 marks)

Adapted from: Science News, 23/04/2011, Vol. 179 Issue 9, p24-25 by Alexandra Witze

Read the article and then answer the questions that follow.

In ancient times, listing the ingredients of the universe was simple: earth, air, fire and water. Today, scientists know that naming all of that, plus everything else familiar in everyday life, leaves out 95 percent of the cosmos's contents'.

From the atoms that make up an astronomer, to the glass and steel of a telescope, to the hot plasma of the stars above -- all ordinary stuff accounts for less than 5 percent of the mass and energy in the universe. "All the visible world that we see around us is just the tip of the iceberg," says Joshua Frieman, an astrophysicist at the University of Chicago and the Fermi National Accelerator L aboratory in Batavia, Ill.

The rest is, quite literally, dark. Nearly one-quarter of the universe's composition is as-yet-unidentified material called dark matter. The remaining 70 percent or so is a mysterious entity known as *dark energy* that pervades all of space, pushing it apart at an ever-faster rate.

A different matter

Dark matter made its debut in 1933, when Swiss astronomer Fritz Zwicky measured the velocities of galaxies in a group known as the Coma cluster and found them moving at different rates than expected. Some unseen and large amount of "dunkle Materie," he proposed in German, must exist, exerting its gravitational effects on the galaxies within the cluster.

Astronom er Vera Rubin confirmed dark matter's existence in the 1970s, after she and colleagues had measured the speeds of stars rotating around the centres of dozens of galaxies. She found that, counter intuitively, stars on the galaxies' outer fringes moved just as rapidly as those closer in - as if Pluto orbited the sun as quickly as Mercury. Rubin's work demonstrated that each galaxy must be embedded in some much larger gravitational scaffold.

The leading candidate for a dark matter particle is the vaguely named "weakly interacting massive particle," or WIMP. Such particles would be "weakly interacting" because they rarely affect ordinary matter, and "massive" because they must exceed the mass of most known particles, possibly weighing in at as much as 1,000 times the mass of the proton. But nobody has yet definitively detected a WIMP, despite decades of experiments designed to spot one.

Mysterious forces

Spotting dark matter may prove to be easier than understanding dark energy, whose mysteries make scientists feel like mental wimps.

Albert Einstein unknowingly ushered dark energy onto the stage in 1917, while modifying his new equations of general relativity. Einstein wondered why gravity didn't make the universe contract in on itself, like a balloon with the air sucked out of it. In 1929, though, E dwin Hubble solved Einstein's problem by reporting that distant galaxies were flying away from each other. The universe, Hubble showed, was expanding. It had been zooming outward ever since the Big Bang gave birth to it. Something funny was going on, giving the cosmos a repulsive push. So in 1998 Michael Turner, a cosmologist at the University of Chicago, dubbed the thing "funny energy" at first, before settling on "dark energy."

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PHYSI	CS 30 STAGE 3
a)	According to the article, all matter that we can observe makes up only 5% of the content of Universe. What constitutes the other 95%? (2 marks)
-	<u>A combination of dark matter (~25%) and dark energy (~70%)</u>
- - b)	What observations did Fritz Zwicky make for him to theorise about the presence of "dunkle
-	<i>He observed that in the Coma cluster of galaxies, the velocities of the</i> <i>galaxies were not as expected. He proposed that there must be other matte</i> <i>exerting gravity on the galaxies in the cluster.</i>
- c)	How did Vera Rubin confirm the existence of dark matter in 1970? (4 marks)
د م +	She observed that outer stars in a galaxy had similar velocities to those close to the centre. This was counter-intuitive as in any normal orbit system, outer satellites have greater periods and correspodingly slower
- -	velocities (eg: Pluto's velocity is much less than ours). She concluded that there must be other matter exerting gravitational effects in a galaxy.
d)	Astronomers have theorized the existence of WIMPS (weakly interacting massive particles) Why do astronomers believe the particles must be massive? (2 marks)
-	Dark matter must have immense mass to exert the observed effect on stars in galaxies. Dark matter particles must therefore have large
-	masses too. Perhaps 1000 times the mass of a proton.
-	

31 e) In 1917, Einstein pondered the fact that the Universe doesn't contract in on itself. Why do you think Einstein would have expected the Universe to contract? (2 marks) Presumably he theorised that all parts of the Universe would be pulled toward the Universe's centre of mass. f) Why is the concept of "Dark Energy" necessary for our current understanding of astronomy? (2 marks) There is evidence that the Universe is expanding (eg: Doppler observations). If the "Big Bang" is over, there must be something else causing it to expand. This "something else" has ben loosely called "Dark Energy". g) Our galaxy, the Milky Way contains approximately 300 billion stars. Assuming that our solar system is 26,000 light years from the centre of the galaxy, estimate the amount of mass that would need to be present at the centre of the galaxy (excluding the possibility of dark matter)for our sun to orbit the centre of the galaxy once every 225 million years. (4 marks) $M = \frac{4\pi^2 r^3}{GT^2}$ $r=2.46 \times 10^{20} m$ $T 7.10 \times 10^{15} s$ $M = \frac{4\pi^2 (2.46 \times 10^{20})^3}{6.67 \times 10^{-11} \times (7.1 \times 10^{15})^2}$ $= 5.56 \times 10^{40} \text{kg}$ (about 28 billion solar masses) End of Questions SEE NEXT PAGE

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	Additional Working Space	
	SEE NEXT PAGE	

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PHYSICS	34	STAGE 3
	END OF EXAMINATION	N